

# Deep Learning Based Distance Estimation of Galaxies

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# Distance to Galaxies

- Hubble-Lemaitre Law

$$v = H_0 d$$

$v$  - recessional velocity

$H_0$  - Hubble's constant

$d$  - proper distance

$$z = \frac{v}{c}$$

$c$  - speed of light

- Distance measured by using **Redshift**

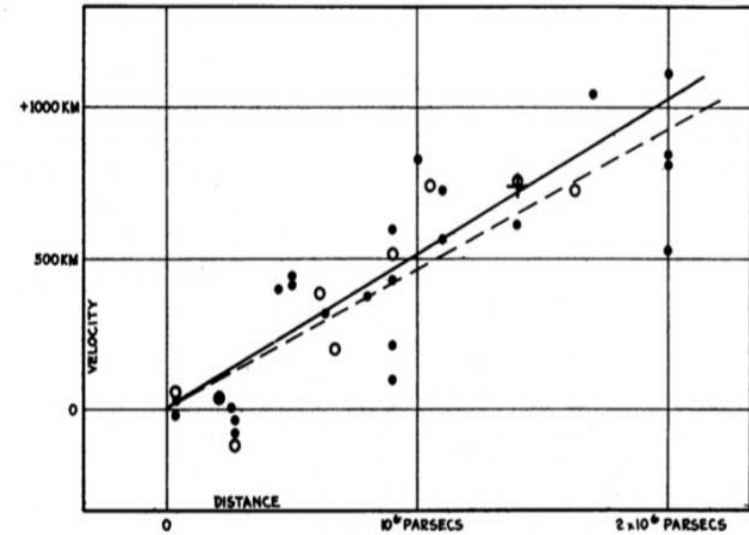


FIGURE 1

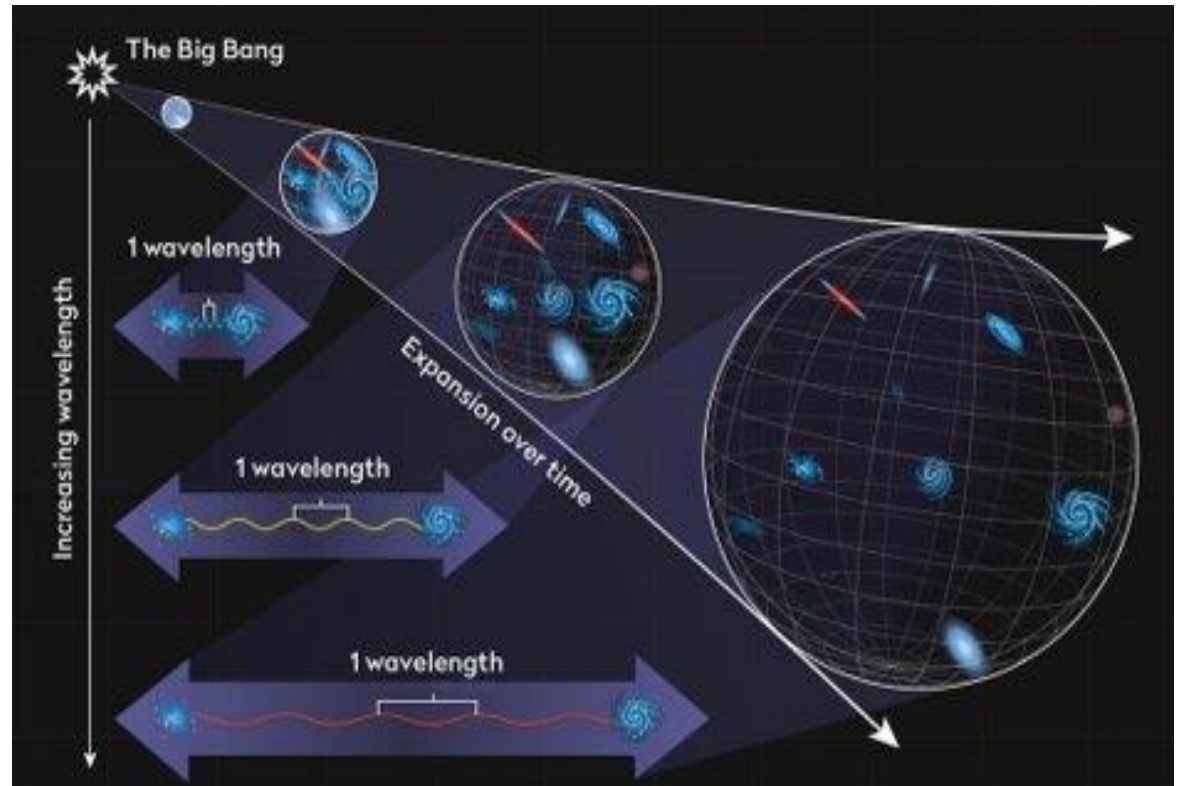
Velocity-Distance Relation among Extra-Galactic Nebulae.

Credit: [Edwin Hubble, Proceedings of the National Academy of Sciences, vol. 15 no. 3, pp.168-173](#)

# Cosmological Expansion and Redshift

- Proper physical distance between a pair of well-separated events is increasing with time.
- The light from galaxies shifted to longer wavelengths - **Redshift (z)**

$$1 + z = \frac{\lambda_{\text{observed}}}{\lambda_{\text{emitted}}}$$



# How to measure Redshift?

## Spectroscopic way

- Measuring the shift in spectral lines
- Spec-z

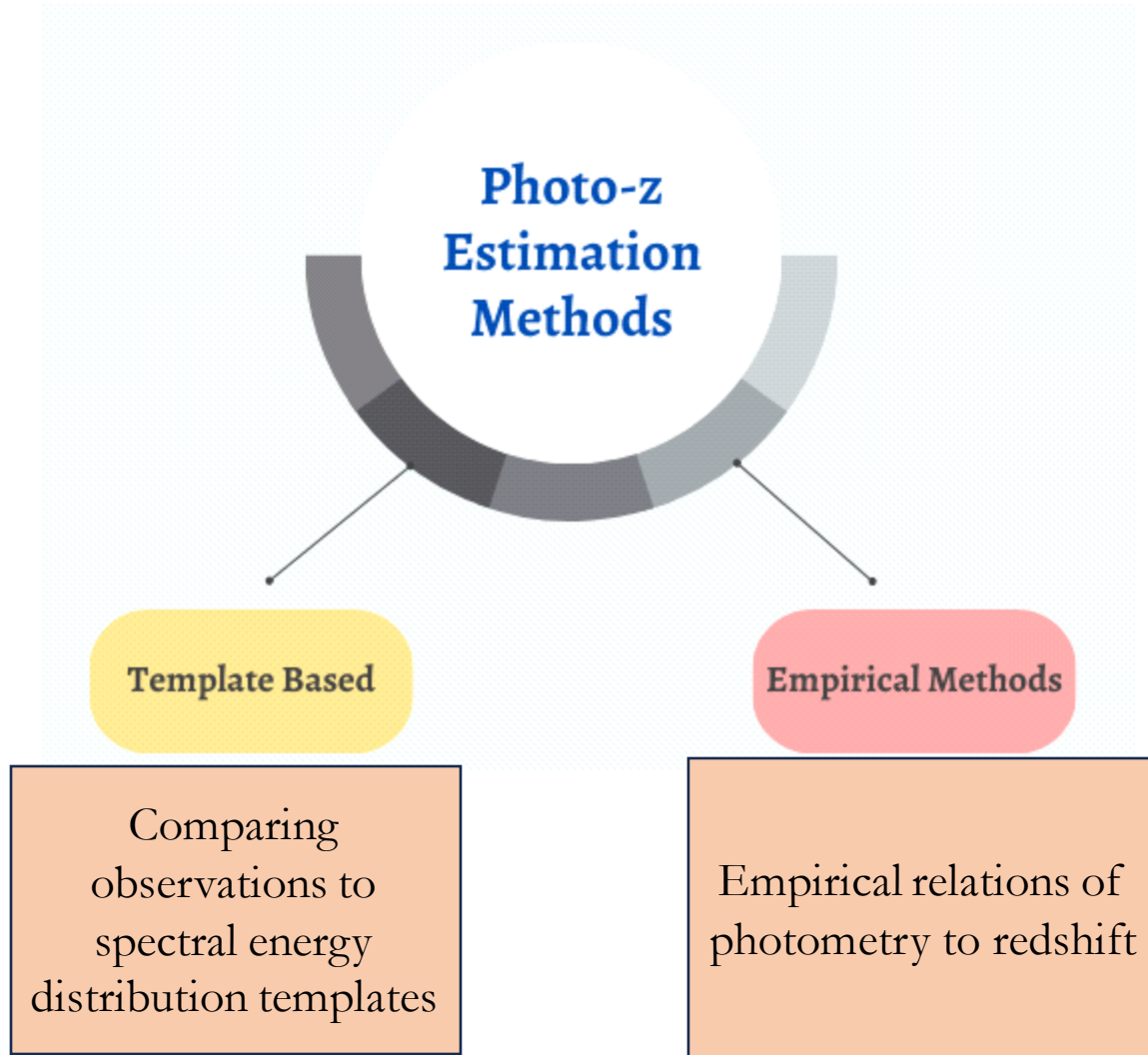
## Photometric way

- Based on observed **photometric quantities**
- Mapping of photometric space into redshift space.
- Photo-z



- The measurable quantities or characteristics of an object's light.
- Derived from the object's intensity or flux measurements.

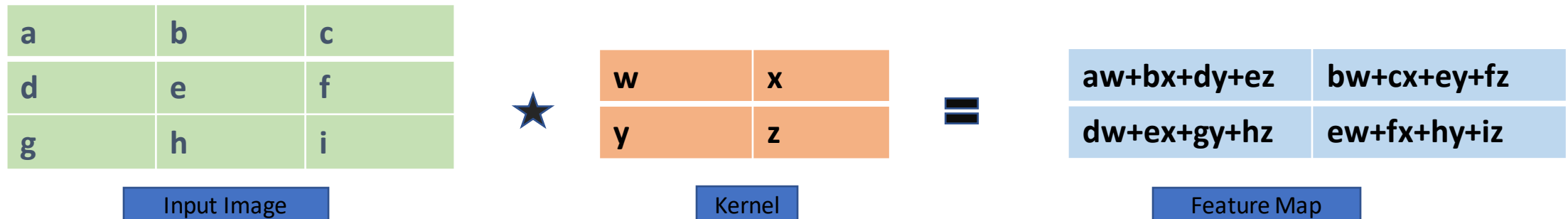
# Photo-z estimation methods



- Deep learning based estimation is an empirical method.
- Convolutional Neural Network (CNN) is used for photo-z estimation of galaxies in Kilo-Degree Survey (KiDS).
- KiDS is a wide-angle image campaign using broad band optical filters (u,g, r and i).

# Convolution

- Small matrix of weights – Kernel/Filter
- Convolved with input data to extract features such as edges, corners etc. of input data.



- Activation function is applied in feature map to introduce non-linearity into the network.

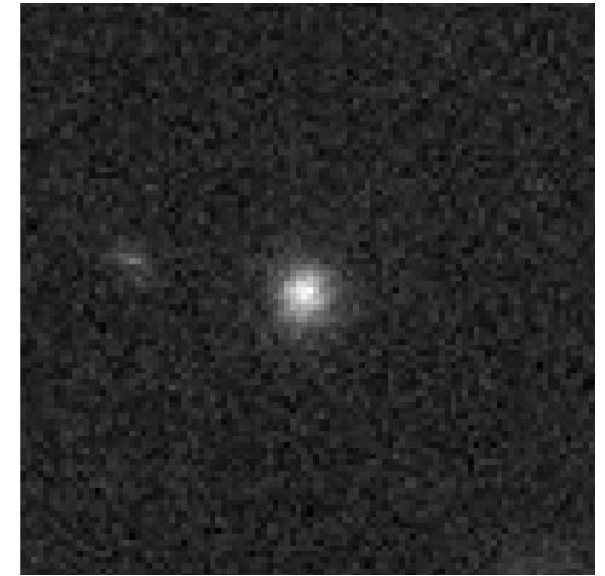
# Pooling and Dense layers

- Pooling layers are used to reduce the dimension of feature map.
- Feature map is a tensor; multi-dimensional array
- It is flattened into 1D array – Input for dense layer
- Performs high-level feature extraction based on low-level features learned by the convolutional layers and pooling layers.

# Input Data



- Input images are galaxy cutouts from KiDS DR4 bright sample.
- Network is trained by **galaxy 4- band images** and their corresponding spectroscopic redshifts (spec-zs).
- **Spec-zs** are obtained from Galaxy And Mass Assembly (GAMA) survey.
- Images are supplement with **9-band magnitudes** (u, g, r, i, Z,Y,J,H and, Ks)
- Cutout size = (36,36,4)
  - Height = 36 pixels
  - Width = 36 pixels
  - Number of bands = 4; (u, g, r and, i)



36x36 cutout of Galaxy image in i-band  
and its rmag=19.37



# Inception

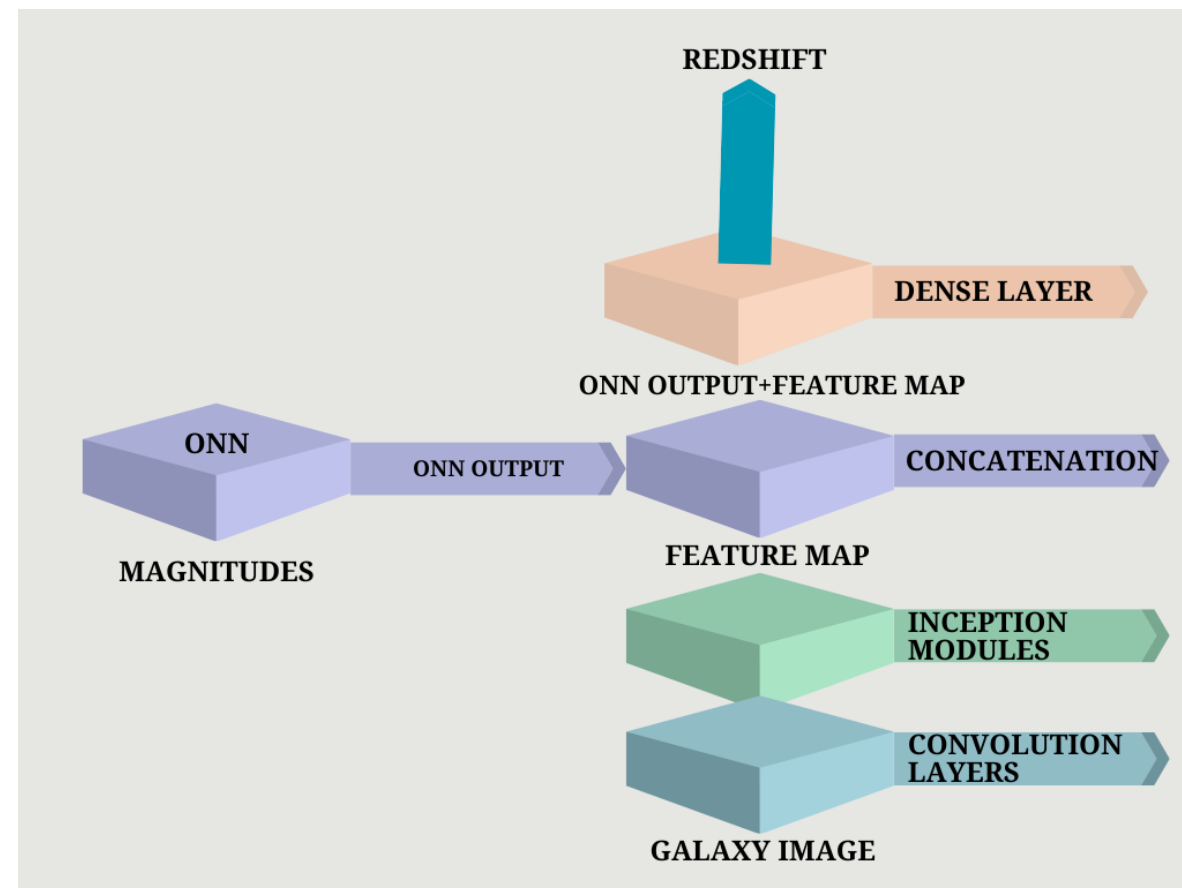
- Inception is a CNN architecture for image recognition.
- Parallel convolution operation

## Inception Module

- Input layer
- 1x1 convolution layer
- 3x3 convolution layer
- 5x5 convolution layer
- Max pooling layer
- Concatenation layer

# Inception based photo-z

- Developed a deep learning model based on Inception.
- Treated as a regression problem.
- This model uses two inputs:
  - Galaxy images
  - Magnitudes of galaxies



- ONN – Ordinary Neural Network
- Concatenation – Combines two outputs

# Training

- Network predicts redshift.
- This predicted redshift is compared with the true redshift by loss function.
- Huber loss function is used.
- It is the combination of Mean Squared Error (MSE) and Mean Absolute Error (MAE).

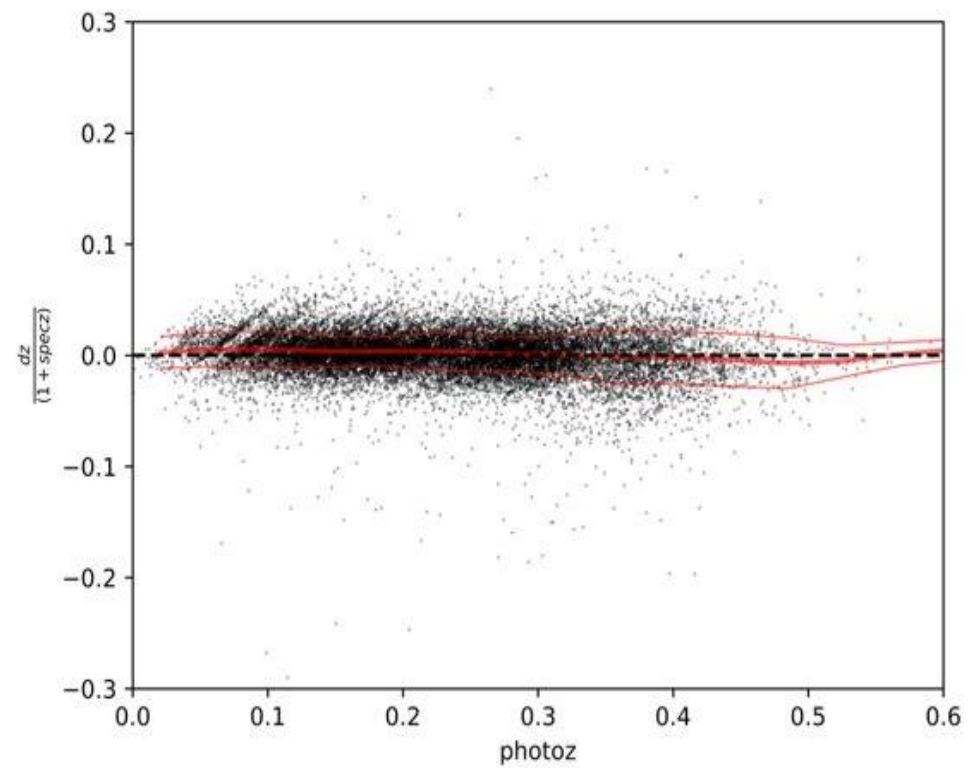
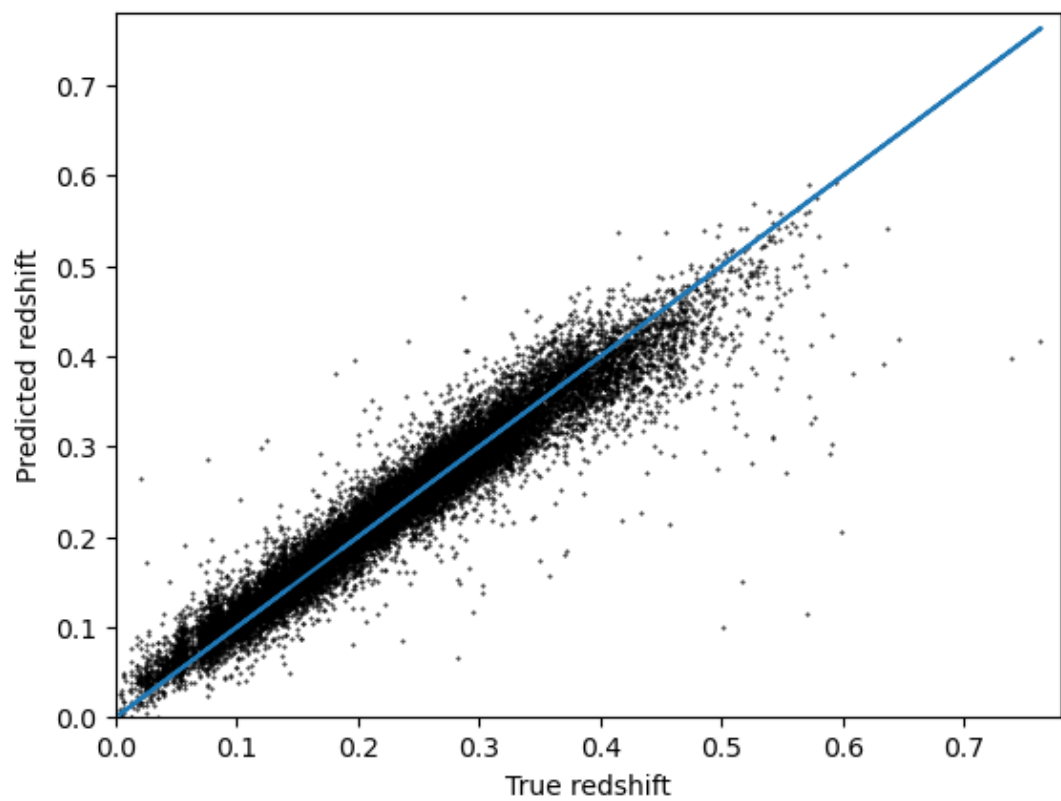
$$L_h = \begin{cases} \frac{1}{2}(e)^2, & |e| \leq \alpha \\ \alpha(|e| - \frac{1}{2}\alpha), & \text{otherwise} \end{cases}$$

- $E = \text{true redshift} - \text{predicted redshift}$
- $\alpha$  is a hyperparameter that determines the transition between MSE and MAE
- During training, the network tries to minimize this loss function by adjusting the weights in kernel.
- Training : Validation : Testing = 70:15:15

# Statistics

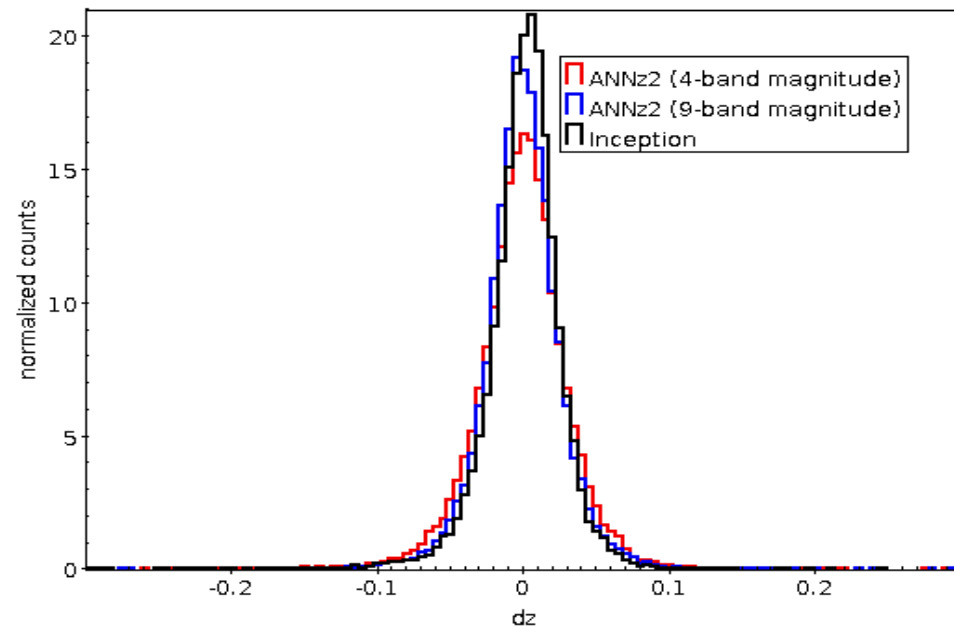
- Bias
  - $dz = \text{Photoz} - \text{Specz}$
- Normalized bias
  - $\text{normdz} = dz / (1 + \text{specz})$
- Standard Deviation of normdz, **SD(normdz)**
- **Median (normdz)**
- Scale Median Absolute Deviation of normdz, **SMAD(normdz)**
  - Where, **SMAD(x) = 1.4826 \* median(|x - median(x)|)**

# Result



# Comparison with ANNz2 Result

- ANNz2 is a photo-z estimation method based on ordinary neural network. (*Bilicki et al. 2018, A&A 616, A69*).



| Method                   | SMAD(normdz) |
|--------------------------|--------------|
| ANNz2 (4-band magnitude) | 0.021        |
| ANNz2 (9-band magnitude) | 0.018        |
| Inception                | 0.016        |

# Future Work

- Apply to KiDS-DR5 bright sample. KiDS-DR5 is the final data release.



Thank You